



# **HerMES: Overview and Coming Results**

**The Herschel Multi-Tiered Extragalactic Survey**

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**Hermes Survey Co-Coordinator with Seb Oliver**

**Through the Infrared Looking Glass  
Pasadena**

**2-5 October 2011**



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Faculty and Researchers, Postdocs, Students



# The HerMES Bibliography

## Science Demonstration Phase Papers (9)

- HerMES: The SPIRE confusion limit Nguyen et al. 2010, A&A 518, L5  
The HerMES SPIRE submillimeter local luminosity function Vaccari et al. A&A 2010, 518, L20  
HerMES: SPIRE galaxy number counts at 250, 350, and 500  $\mu\text{m}$  Oliver et al. 2010, A&A 518, L21  
Halo occupation number and bias properties of dusty galaxies from clustering measurements, Cooray et al. 2010, A&A 518, L22  
First results from HerMES on the evolution of the submillimetre luminosity function, Eales et al. A&A 2010, 518, L23  
Herschel unveils a puzzling uniformity of distant dusty galaxies, Elbaz et al. A&A 2010, 518, L29  
The far-infrared/radio correlation as probed by Herschel, Ivison et al. A&A 2010, 518, L31  
HerMES: The submillimeter spectral energy distributions of Herschel/SPIRE-detected galaxies, Schulz et al. 2010, A&A 518, L32  
HerMES: Far infrared properties of known AGN in the HerMES fields, Hatziminaoglou et al. 2010, A&A 518, L33

## Papers since SDP (23)

- HerMES: Lyman-Break Galaxies Individually Detected in the HerMES Fields, Smail et al. 2011, MNRAS, 413, 1777.  
Modeling of the HerMES J105751.1+573027 Lensed Galaxy, Amblard et al. 2011, Nature, 470, 510.  
Dynamical Structure of the Molecular interstellar Medium in the HerMES J105751.1+573027 Lensed Galaxy, Amblard et al. 2011, MNRAS, 413, 1777.  
Redshift Determination and CO Line Excitation Modeling of the HerMES J105751.1+573027 Lensed Galaxy, Amblard et al. 2011, MNRAS, 413, 1777.  
Discovery of a Multiply-Lensed Submillimeter Galaxy in the HerMES J105751.1+573027 Field, Amblard et al. 2011, MNRAS, 413, 1777.  
HerMES: Cosmic Magnification of Sub-mm Galaxies, Amblard et al. 2011, MNRAS, 413, 1777.  
HerMES: SPIRE Emission from Radio-Selected Active Galactic Nuclei, Seymour et al., 2011, MNRAS, 413, 1777.  
Sub-millimetre Galaxies Reside in Dark Matter Halos with  $M > 3e11$  Solar Masses, Amblard et al, 2011, Nature, 470, 510.  
Measures of star formation rates from FIR and UV emissions of galaxies in HerMES fields, Buat et al. 2010, MNRAS 409, L1.  
HerMES: Far-Infrared observations of Lyman Break Galaxies, Rigopoulou 2010, MNRAS 409, L7  
The Deep SPIRE HerMES Survey: Secure SEDs and their Astrophysical Indications, Brisbin et al. 2010, MNRAS 409, 66.  
Herschel-SPIRE, Far-Infrared Properties of Millimetre-Bright and -Faint Radio Galaxies, Chapman et al. 2010, MNRAS 409.  
HerMES : SPIRE detection of high redshift massive compact galaxies in GOODS-N field, Cava et al. 2010, MNRAS 409, L19.  
Cold dust and young starbursts: SEDs of SPIRE sources from HerMES, Rowan-Robinson et al. 2010, MNRAS 409, 2.  
Herschel reveals a  $T_{\text{dust}}$  - unbiased selection of  $z \approx 2$  ULIRGs, Magdis et al. 2010 MNRAS 409, 22.  
HerMES: Source Extraction and Cross-IDs in Confusion-Limited SPIRE Images, Roseboom et al. 2010, MNRAS 409, 48.  
Evolution of Dust Temperature of Galaxies through Cosmic Time as seen by Herschel, Hwang et al. 2010 MNRAS 409, 75.  
HerMES: SPIRE Science Demonstration Phase Maps, Levenson et al. 2010 MNRAS 409, 83.  
HerMES: Deep Galaxy Number Counts from P(D) of SPIRE SDP Observations, Glenn et al. 2010, MNRAS 409, 109.  
HerMES: The X-Ray Infrared Correlation of Star-Forming Galaxies at  $z \sim 1$ , Symeonidis et al. 2011, MNRAS, in press.  
HerMES: Point Source Catalogs from Deep Herschel-SPIRE Observations, Smith et al. 2011, arXiv 1109.5186  
The Herschel Multi-Tiered Extragalactic Survey: SPIRE-mm Photometric Redshifts, Roseboom et al. 2011, arXiv 1109.2887  
The Herschel Multi-Tiered Extragalactic Survey: HerMES, Oliver et al. 2011, MNRAS, in press.

**HerMES DR1 key papers**  
**18 central and cohesive**  
**publications coming soon!**

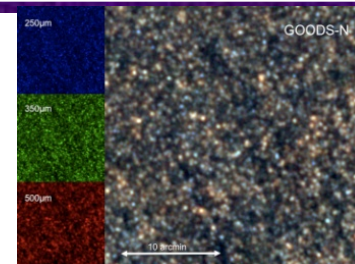


# Herschel Large High-z Surveys

## HerMES: Herschel Multi-tiered Extragalactic Survey

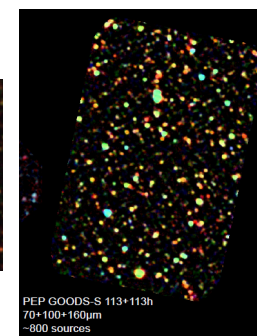
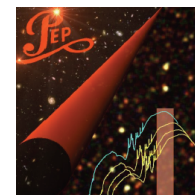


- PACS + SPIRE
- 70 sq deg from 20'×20' to 3.6°×3.6° (900 hours) + 12 clusters
- Bolometric luminosities of galaxies, cosmic SFH
- Wedding cake to probe range of luminosities and environments



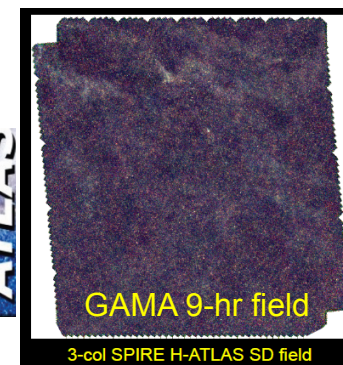
## PEP: PACS Evolutionary Probe

- PACS only
- 2.7 sq deg from 10'×15' to 85'×85' (655 hours) + 10 clusters
- Resolve CFIRB;  $L_{\text{FIR}}$  & SFRs



## H-ATLAS: Herschel-Astrophysical Terahertz Large Area Survey

- PACS + SPIRE
- 550 sq deg (600 hours)
- Large-scale structure, AGN, rare objects
- Expect ~500,000 detections to  $z \sim 3$ , majority at 250 & 350  $\mu\text{m}$



## H-GOODS: Herschel-Great Observatories Origins Deep Survey

- PACS very deep imaging of the GOODS Field (330 hours)
- SPIRE deep imaging of the GOODS Field (30 hours)





# HerMES = SPIRE GT Program

## Spectral and Photometric Imaging Receiver

### Photometer

- 250, 350, 500  $\mu\text{m}$  (simultaneous)
- 4 x 8 arcminute field of view
- Diffraction limited beams (18, 25, 36")

*Fast scan mapping at long wavelengths*

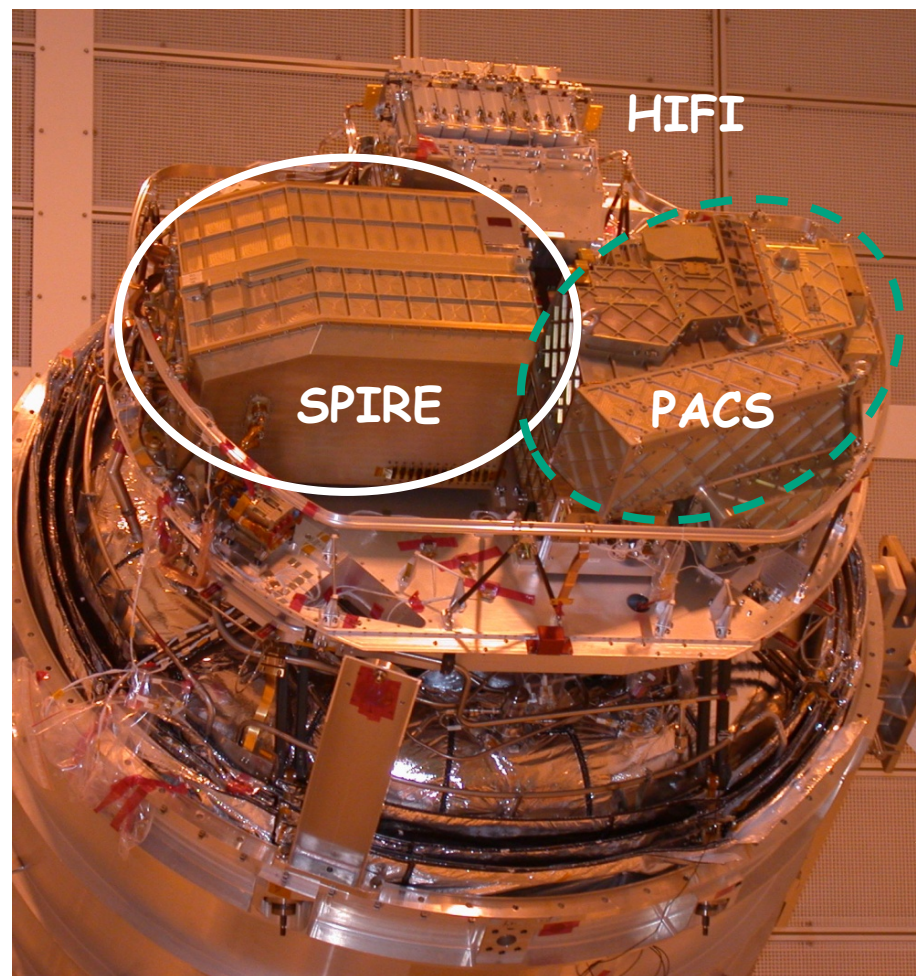
### Imaging FTS

- 200 - 670  $\mu\text{m}$
- 2.6 arcminute field of view
- $\Delta\nu = 1.2$  GHz high resolution mode
- $\Delta\nu = 25$  GHz low resolution mode

*Wide instantaneous bandwidth, map making*

### Design Principles

- Sensitivity limited by thermal emission from the telescope
- $^3\text{He}$  cooled detector arrays (0.3 K)
- Feedhorn-coupled spider-web bolometers
- Minimal use of mechanisms
  - Beam steering mirror; FTS mirror drive
- Optimized for scan-mapped surveys

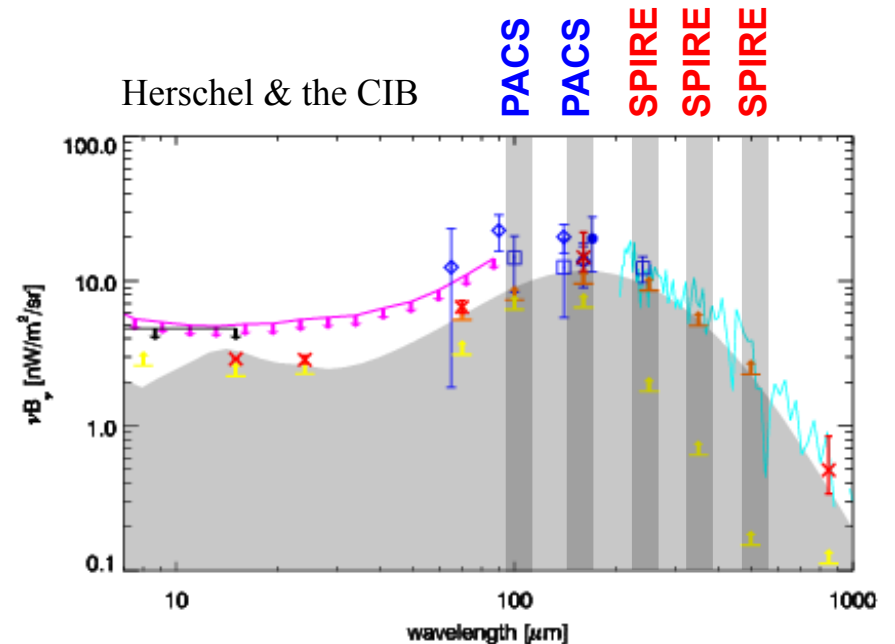
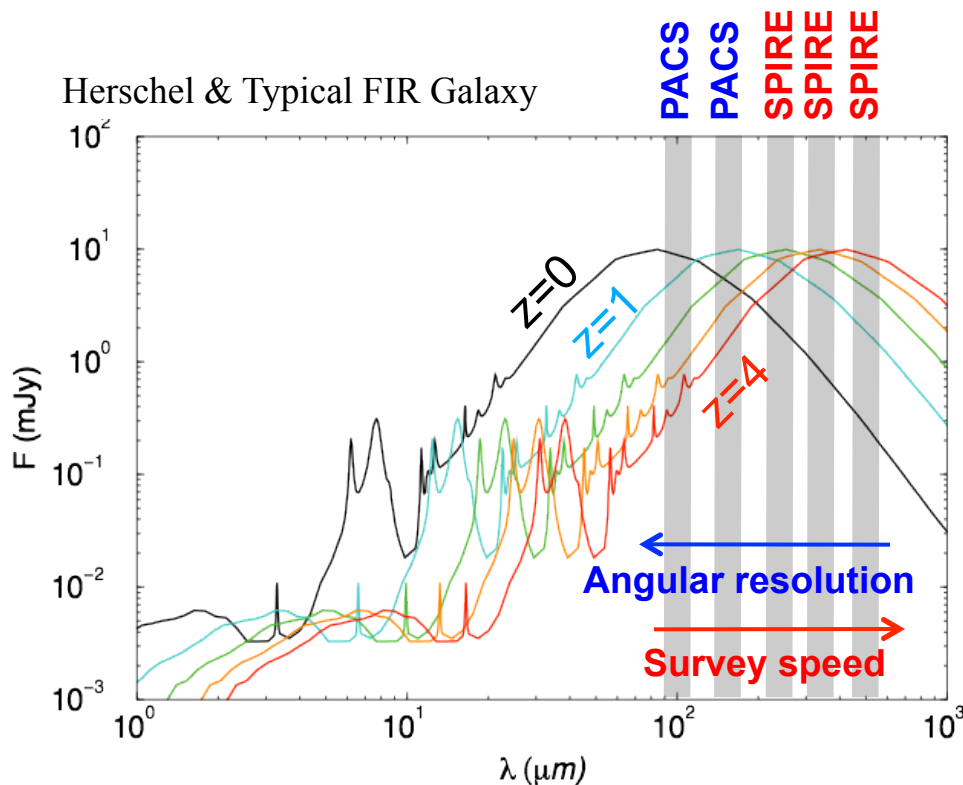




# HerMES Science Motivation

## What is the history of Far-IR galaxies?

- How do they assemble and evolve over time?
- Where have luminous FIR systems gone today?
- How do FIR galaxies relate to dark matter?
- What is the role of dust in star formation?
- What is the connection between dusty star formation and AGNs?



## Herschel Extragalactic Surveys

- Observe at SED peak
- Bolometric far-IR luminosities
- Large and uniform samples



# HerMES Survey Design Principles

## Wedding Cake Design

- Probe a wide range of the luminosity function
- Deep fields for sub-confusion studies
- Wide fields for rare objects and fluctuations

## Use Best Ancillary Fields Available

- Fields with Spitzer, Radio, Optical, NIR, X-ray, etc data

## Do What Herschel Does Best

- SPIRE excels at large maps near confusion limit
- PACS best at small deep maps
- Collaborate with PEP for PACS data
- Use parallel mode where possible



# HerMES: Wedding Cake Survey

## Clusters

L1 0.11  $\square^\circ$

L2 0.36  $\square^\circ$

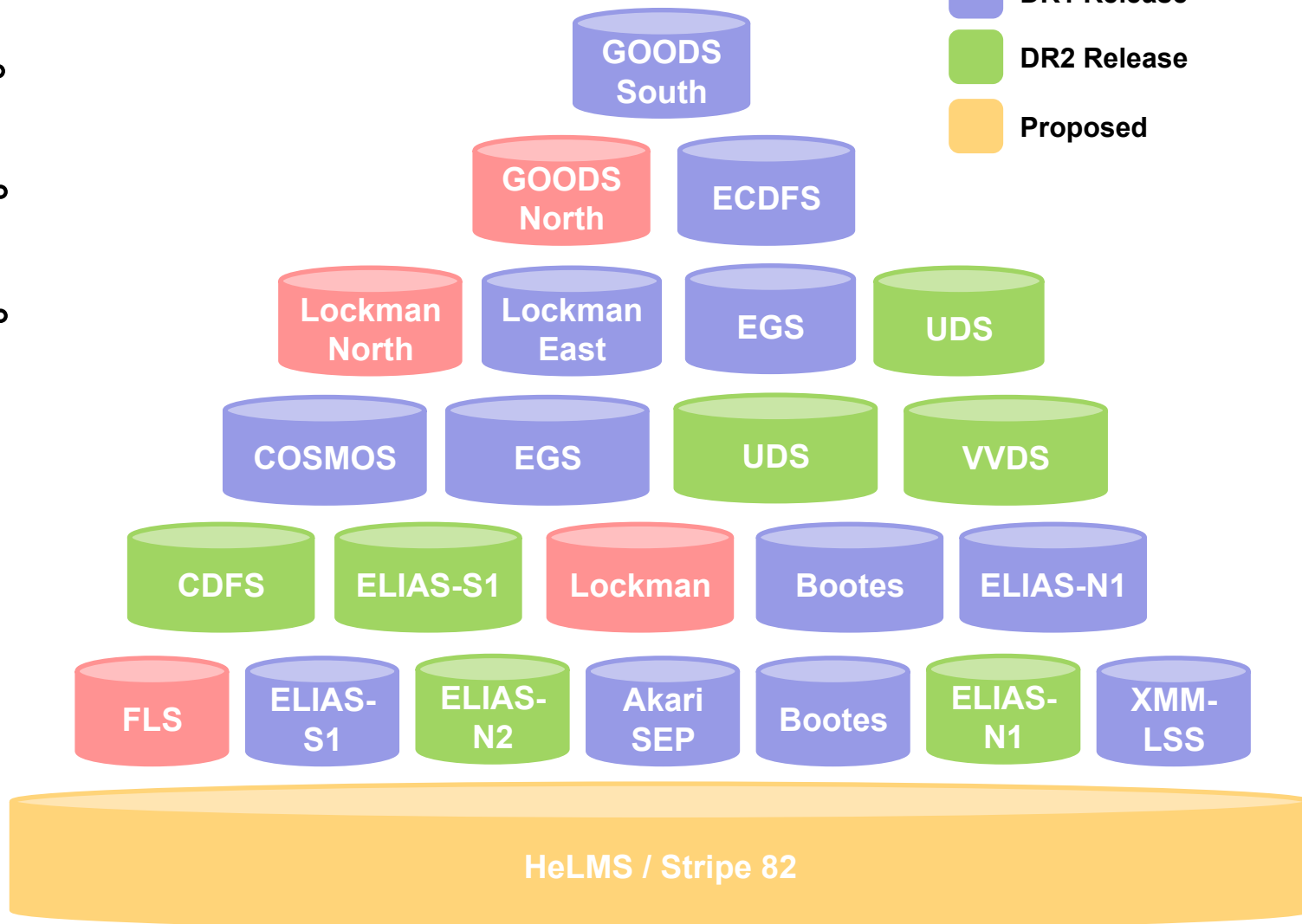
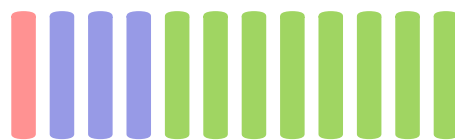
L3 1.25  $\square^\circ$

L4 4  $\square^\circ$

L5 30  $\square^\circ$

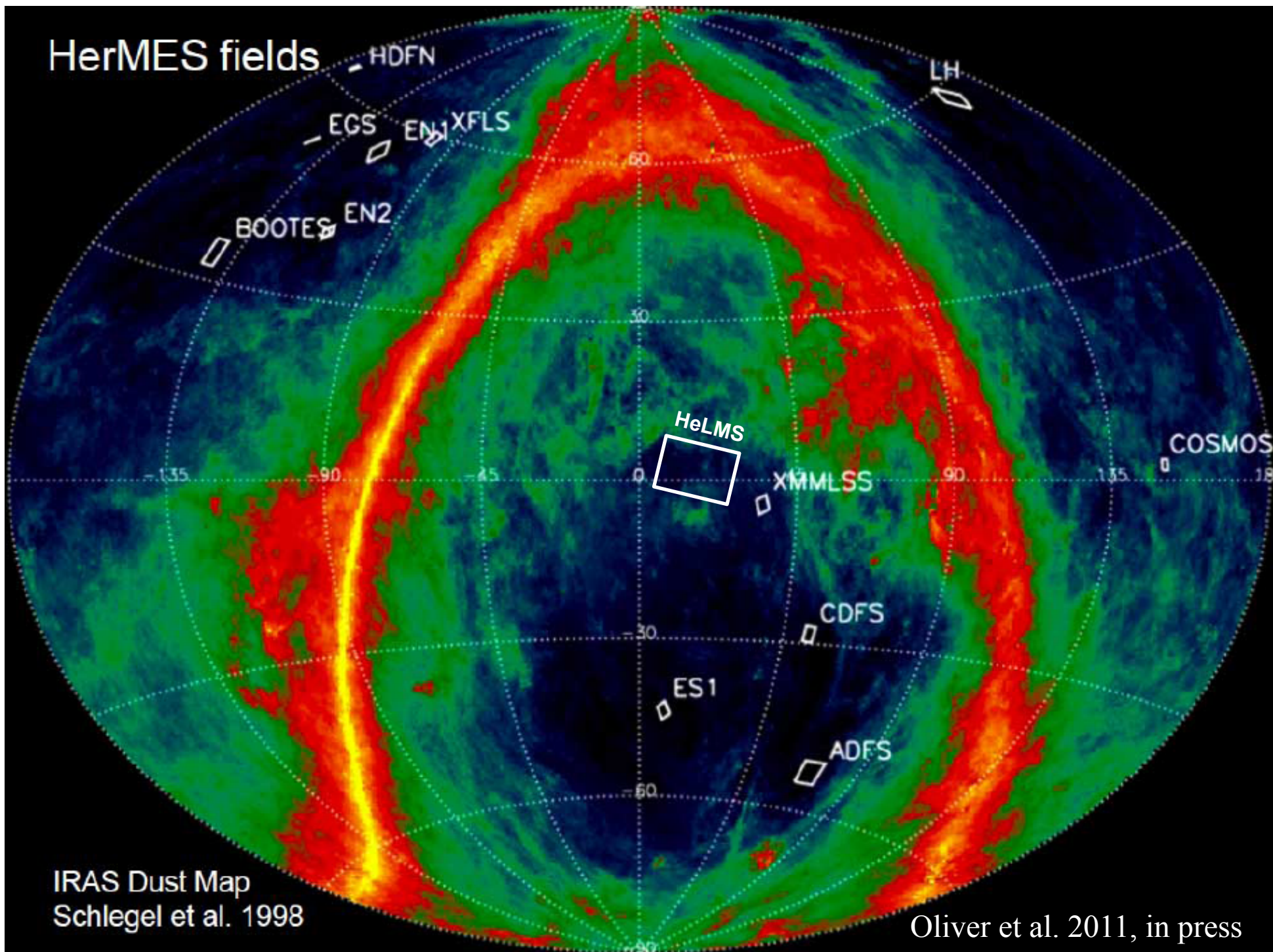
L6 40  $\square^\circ$

L7 270  $\square^\circ$





# HerMES fields



IRAS Dust Map  
Schlegel et al. 1998

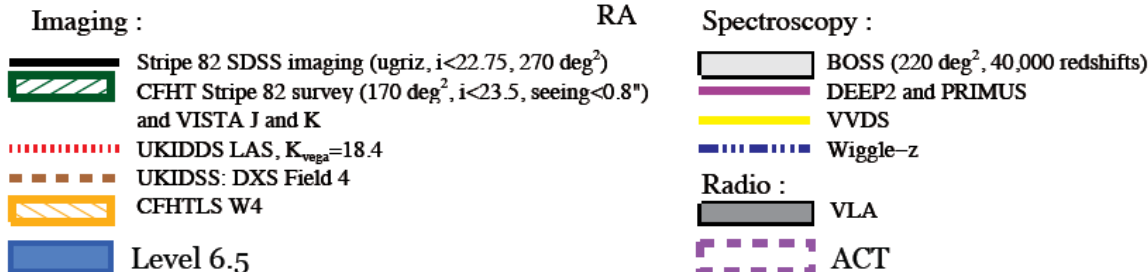
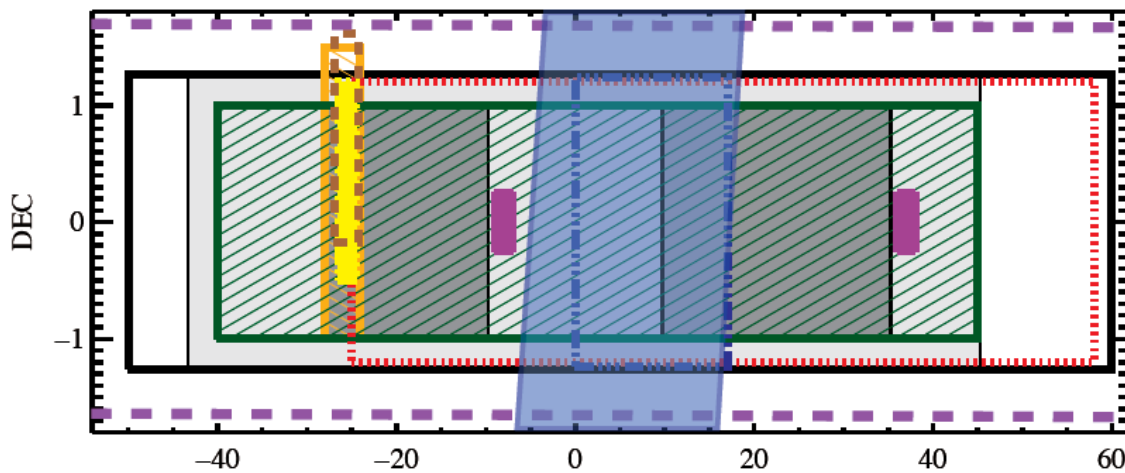
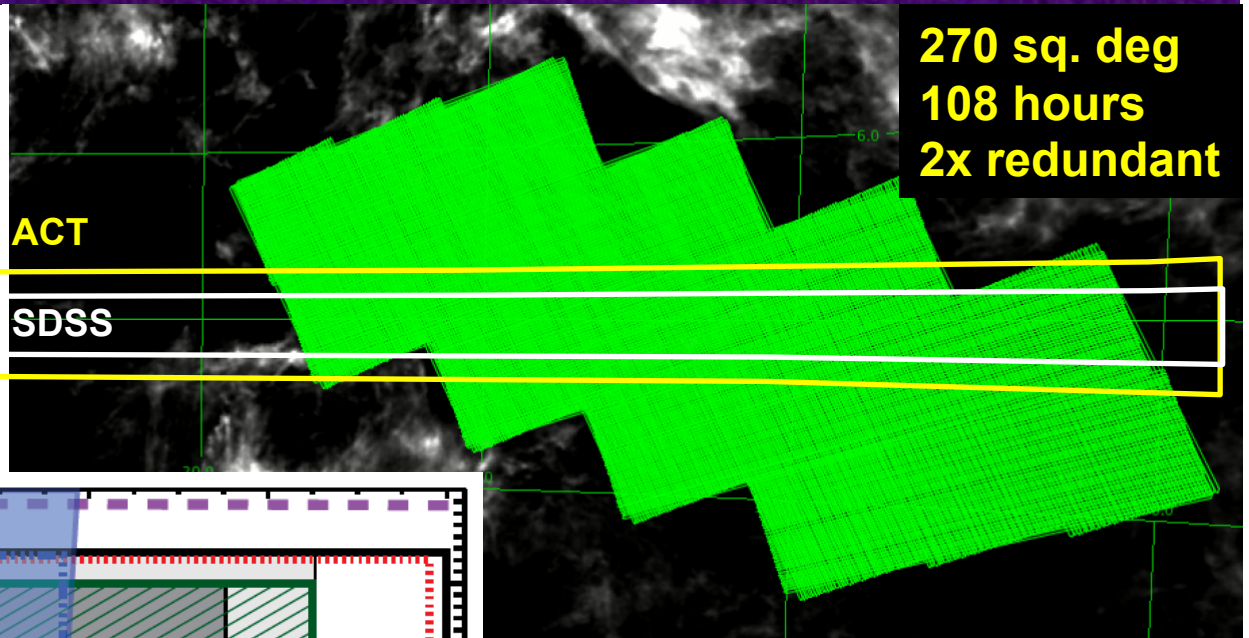
Oliver et al. 2011, in press



# HeLMS: A Cross-Linked Shallow Survey

## Ancillary Data Science

- BOSS  $\sim 10^6$  LRGs,  $z = 2$  quasars
- Hetdex Ly- $\alpha$  galaxies out to  $z \sim 4$
- ACT mm-wave CMB survey
- SDSS, UKIDSS, VLA, Wiggle-z
- Great legacy value



## Luminous Sources

- $\sim 25,000$  galaxies at  $5\sigma$  at  $250 \text{ um}$
- $\sim 250$  lensed galaxies

## Background Fluctuations:

- Extends galaxy x galaxy coverage to  $\ell \sim 100$
- CMB anisotropies x FIR galaxies
- CMB lensing x FIR galaxies

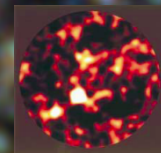
# Selected HerMES Science Highlights And Coming Key Papers

250 $\mu$ m

350 $\mu$ m

500 $\mu$ m

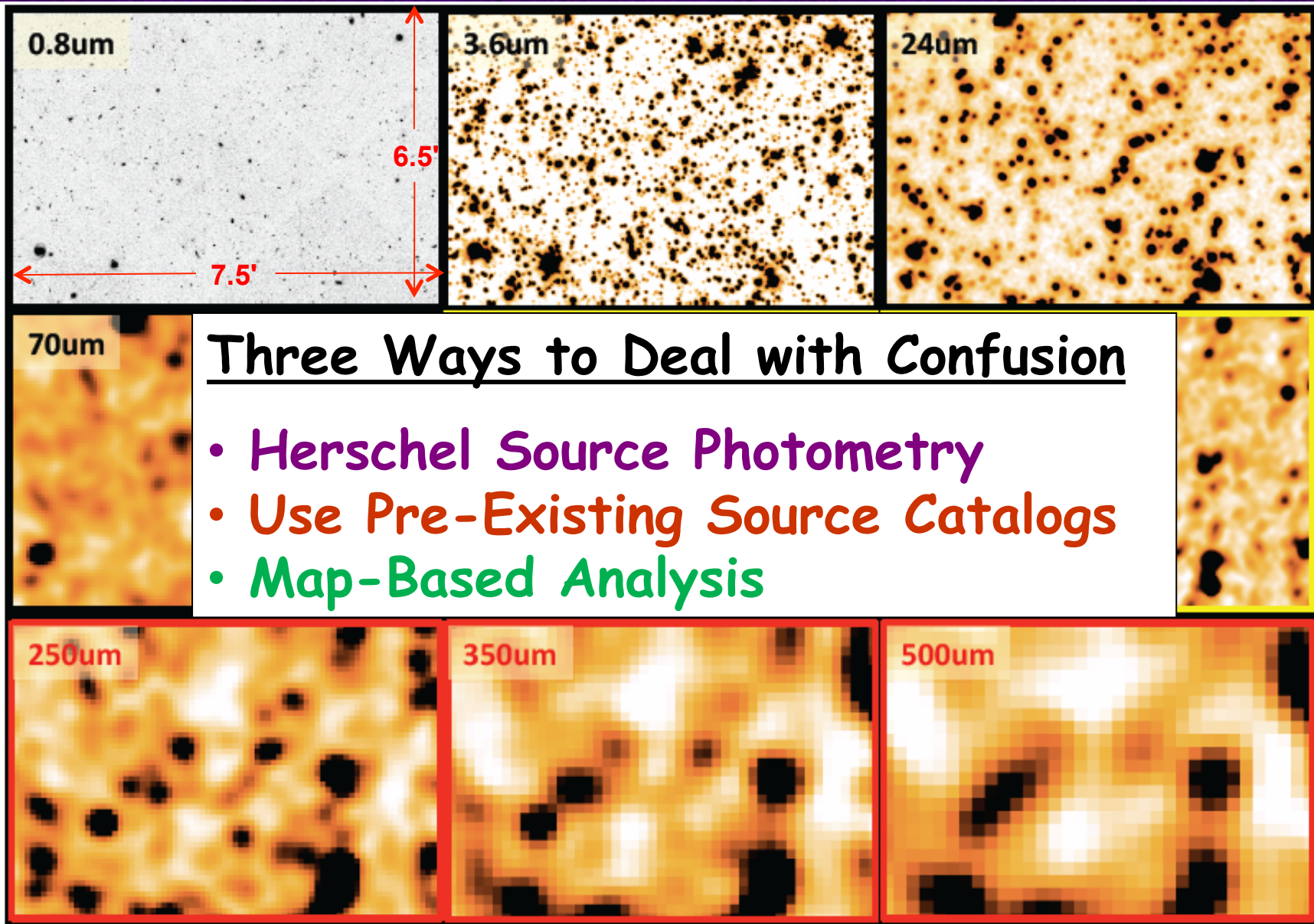
10 arcmin

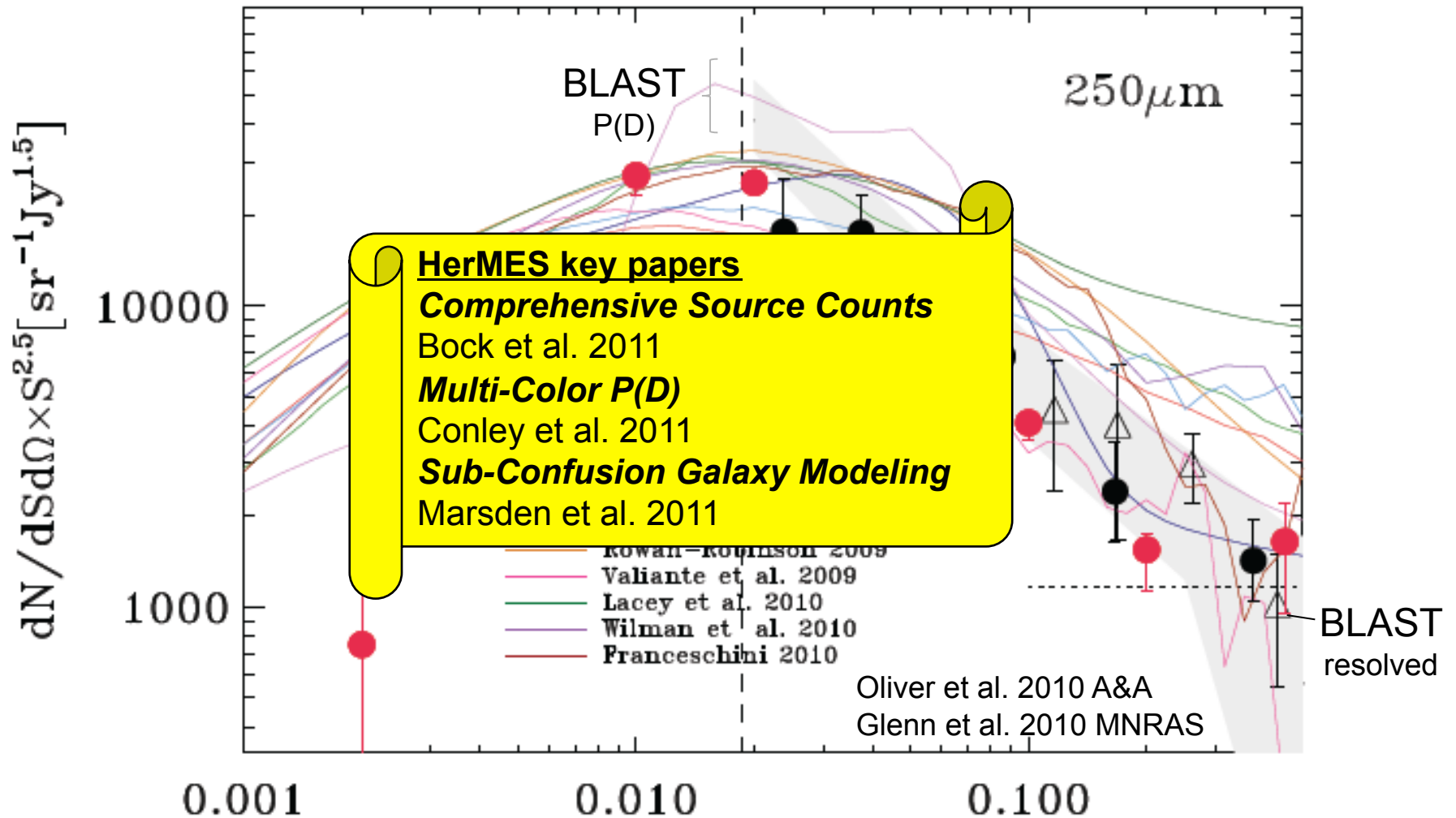


SCUBA HDF

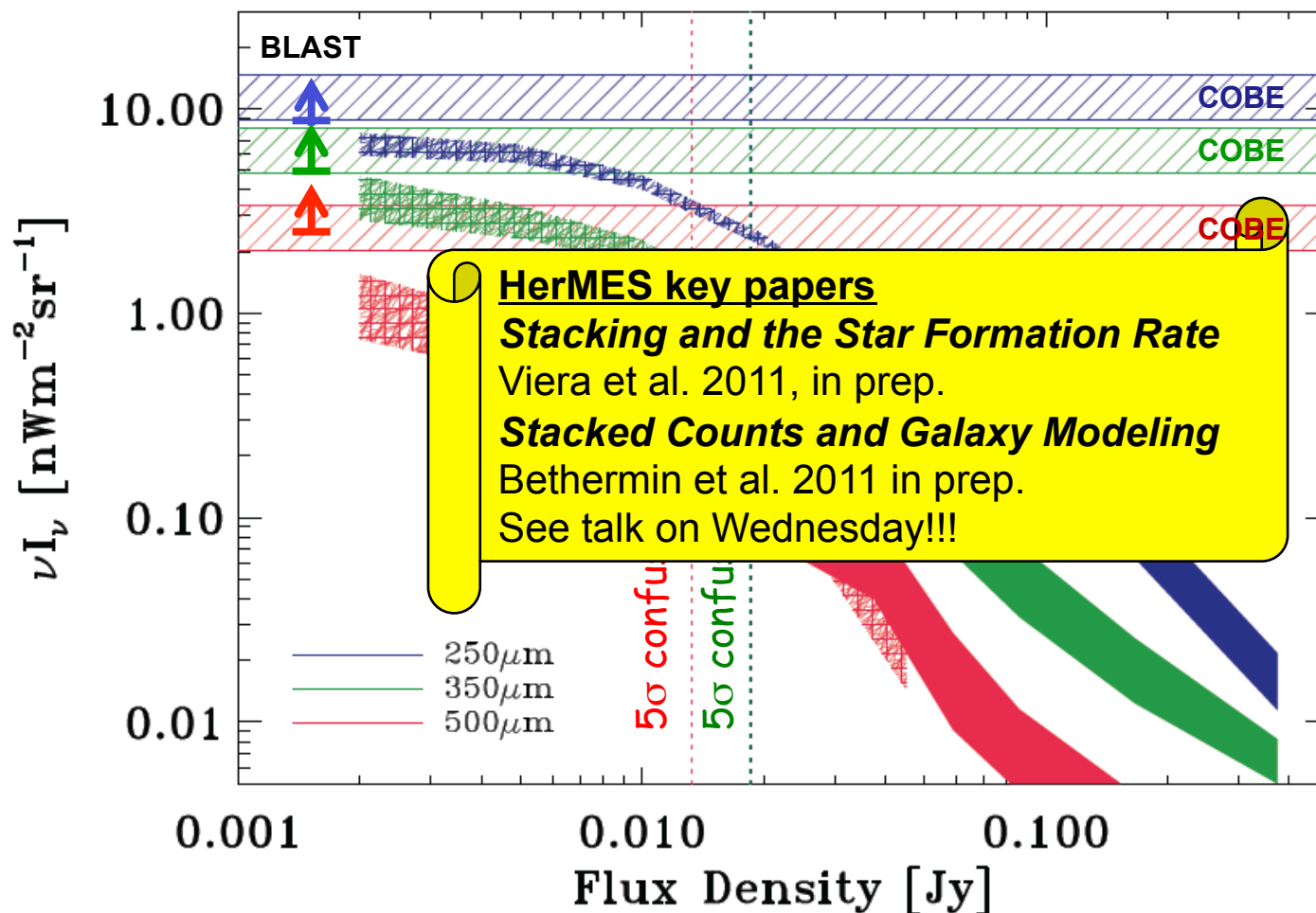
GOODS-N

# The Confusion Challenge





- Number counts of bright galaxies (ULIRGS+) over-predicted by models
- Bright-end counts are steeper than models generically



- **Source Counts**  
 250, 350, 500 μm  
 15%, 10%, 6%

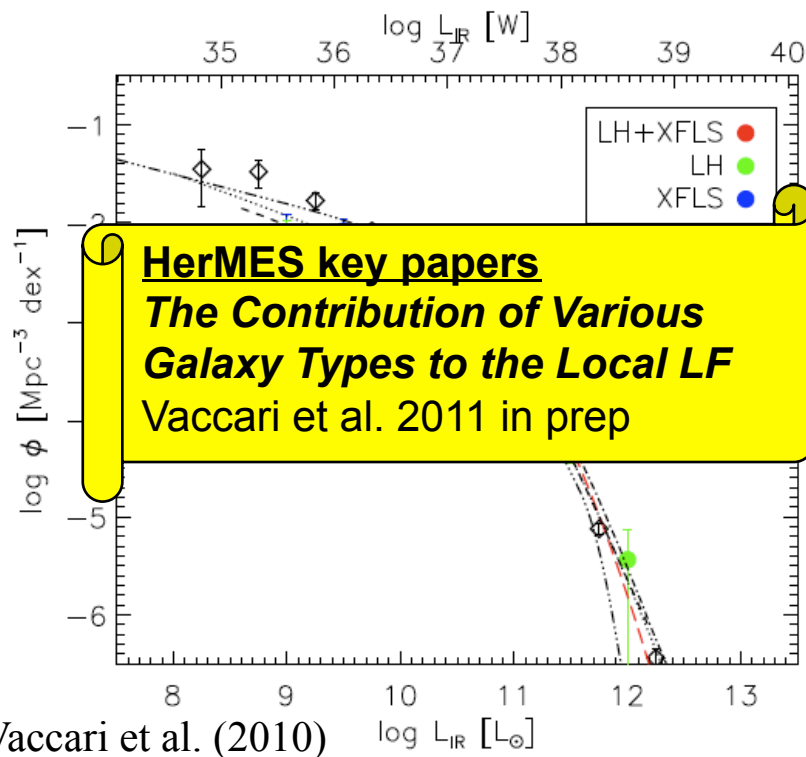
- **P(D)**  
 250, 350, 500 μm  
 65%, 60%, 45%

- **Stacking in prep**  
 With BLAST:  
 250, 350, 500 μm  
 80%, 80%, 85%

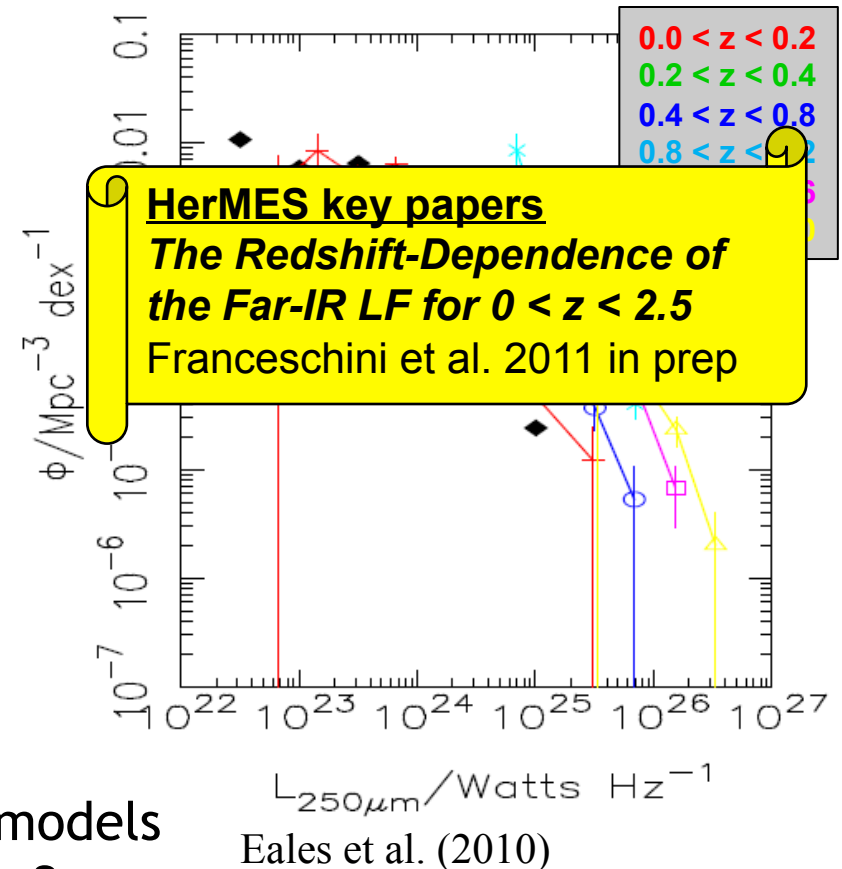
Marsden et al. 2010

Of course: The remainder are the most interesting sources!  
 E.g.  $z > 3$  galaxy populations

## Local Luminosity Function

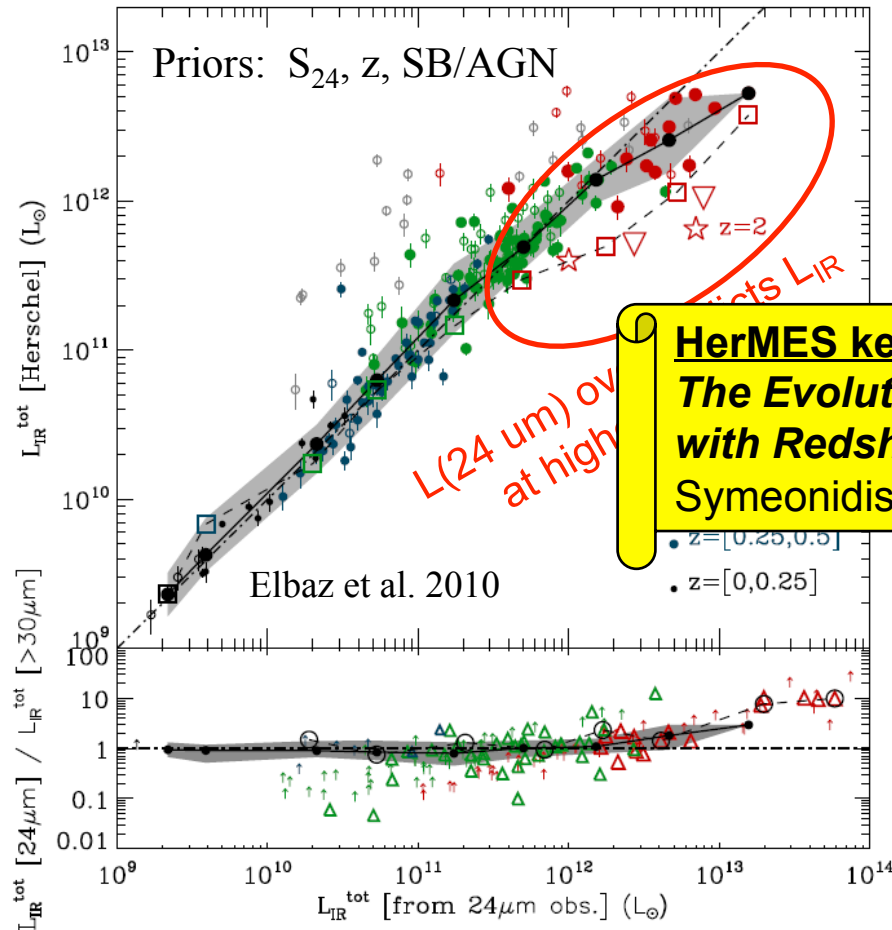


## HerMES Rest-Frame 250 $\mu\text{m}$ LF

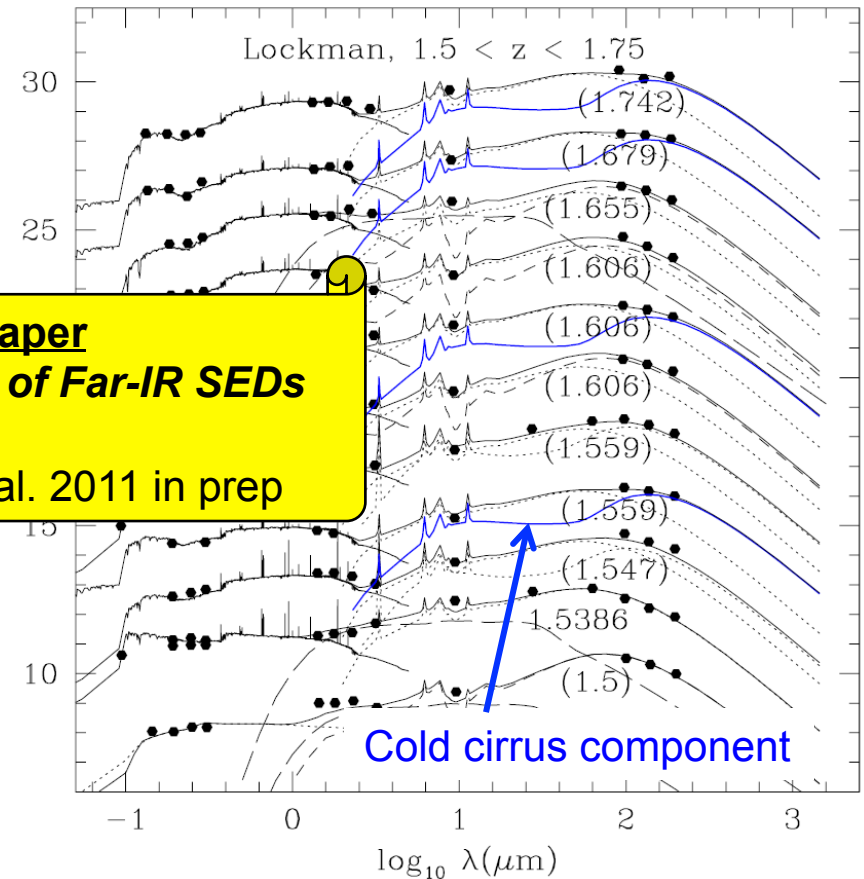


- Local sub-mm galaxy LF slightly above models
- Luminosity function increases out to  $z \sim 2$
- Is it flattening out at  $z > 1$ ?
- Next: better statistics from bigger samples

## $L_{\text{IR}}$ for Starbursts and AGNs



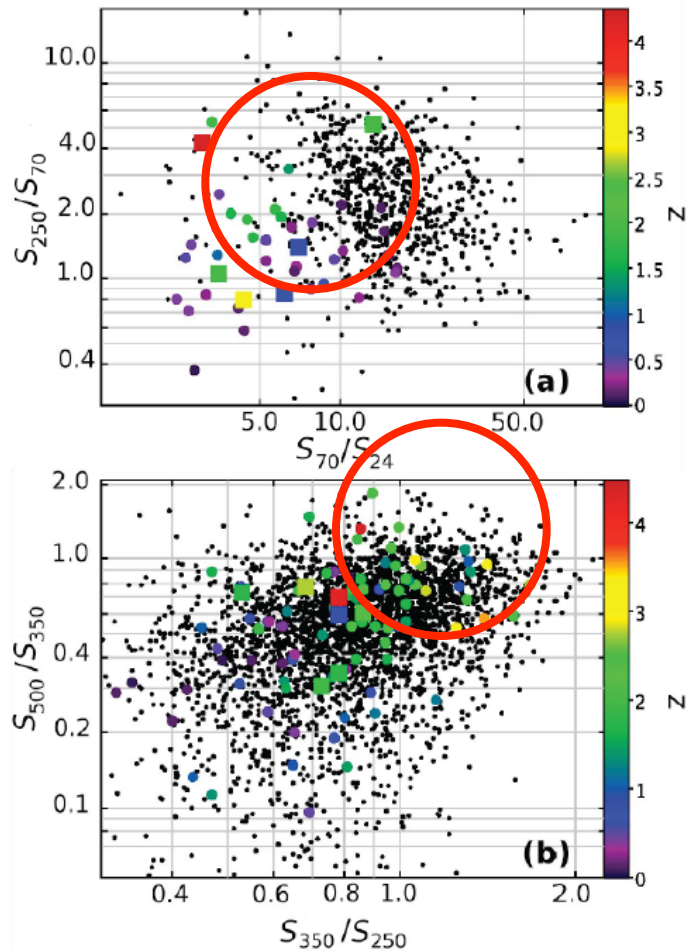
## Multi-Wavelength SED Fits



Rowan-Robinson et al. 2010

- *Herschel* provides a direct measure of bolometric luminosity and SFR
- $L_{\text{FIR}}$  and SFR predicted from  $\lambda \leq 24 \mu\text{m}$  observations are inadequate
- ~Half the SEDs require lower temperature dust component (10 - 20 K)

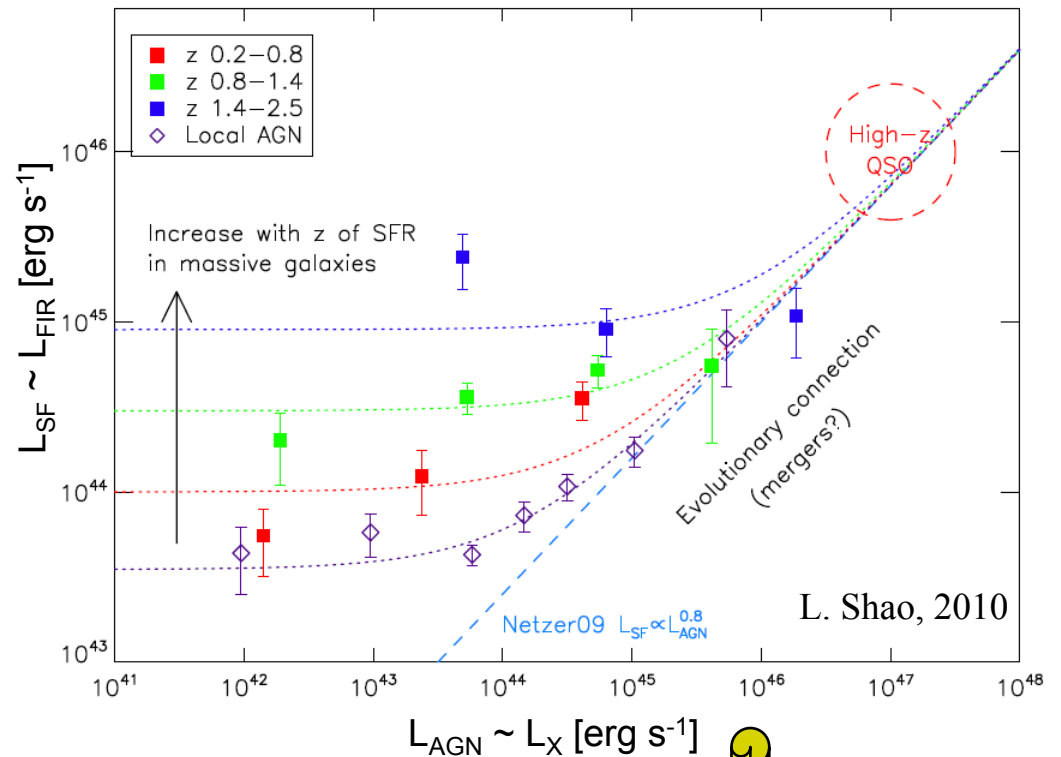




**Distinct  $S_{70}/S_{24}$  colors, but not  $S_{250}/S_{350}$**   
 - FIR emission due to star formation

E. Hatziminaoglou, 2010

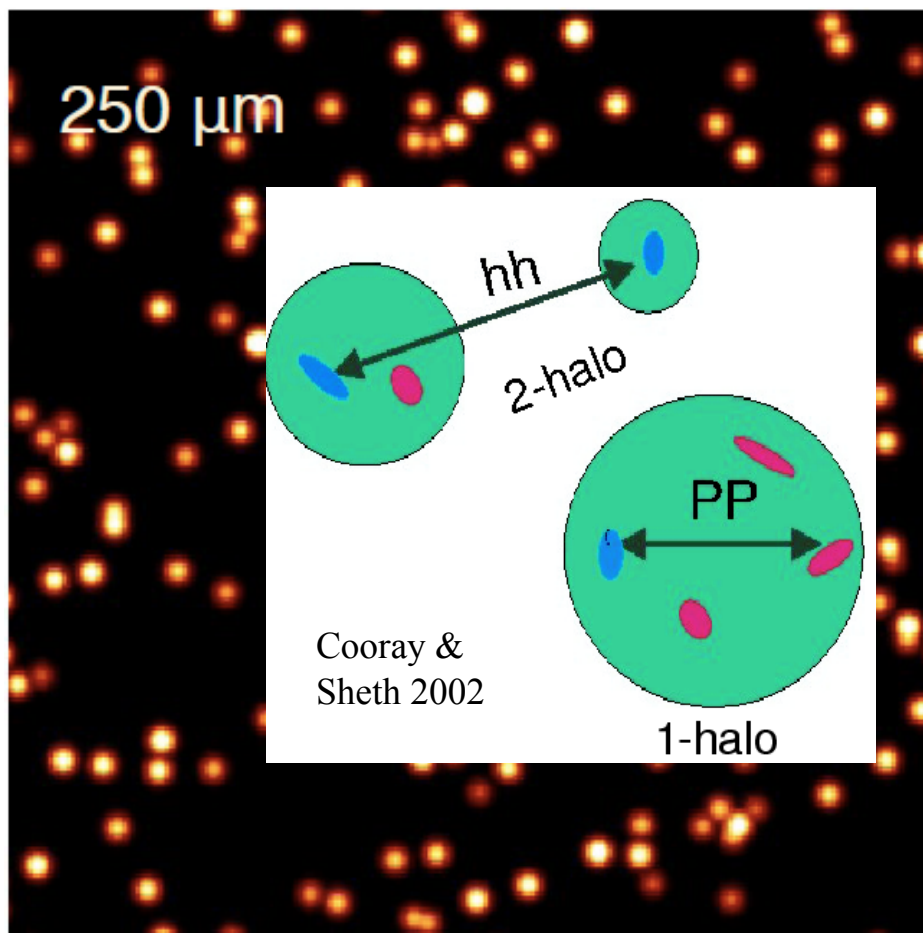
## Two Modes of Host & AGN Evolution?



**HerMES key paper**  
**How do AGNs relate to SFR traced in SPIRE bands?**  
 Page et al. 2011 in prep

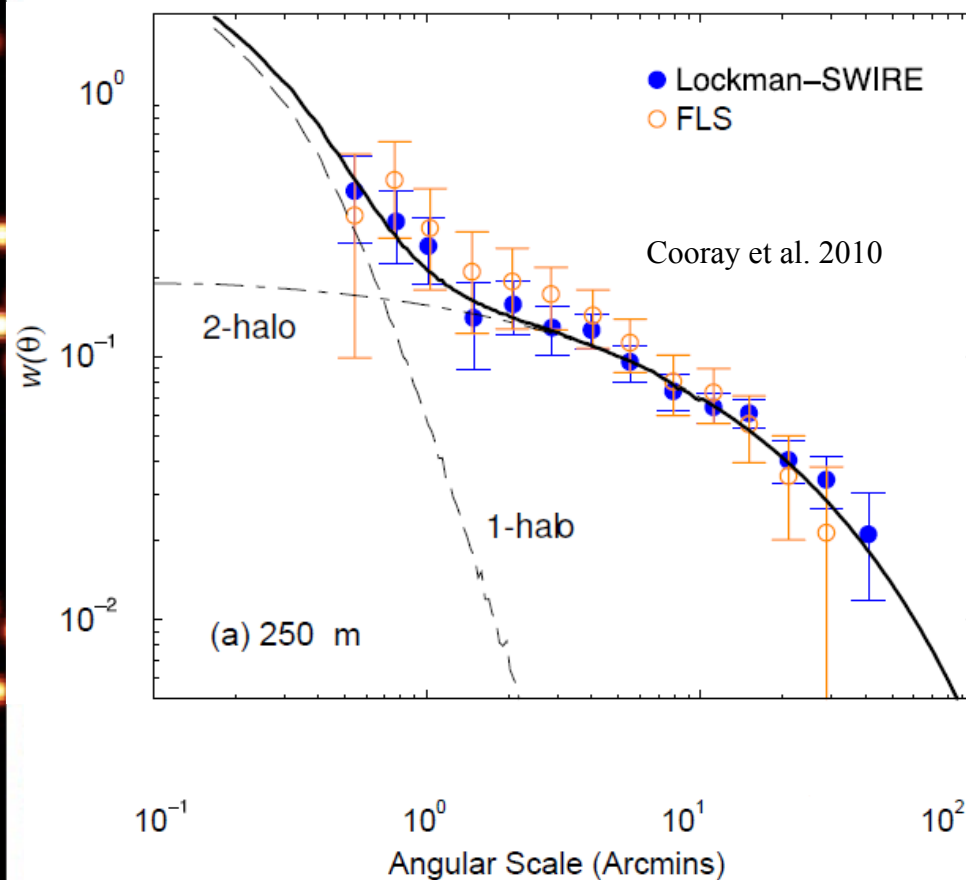
**High L<sub>AGN</sub> related to L<sub>SF</sub>**  
 - Trend is weaker  
 - Luminosities coupled by mergers?

$L_{SF}$  w/  $z$  galaxies



$S > 20 \text{ mJy} : 1,200/\text{deg}^2$

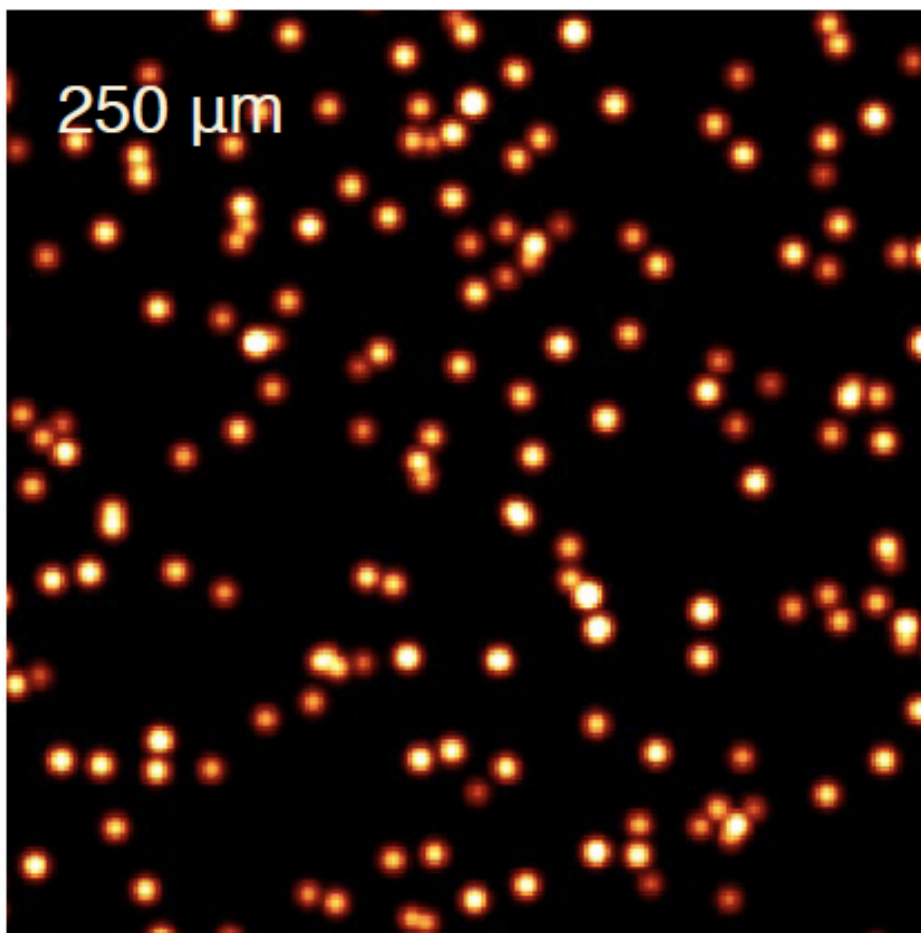
Angular Correlation of Detected Galaxies



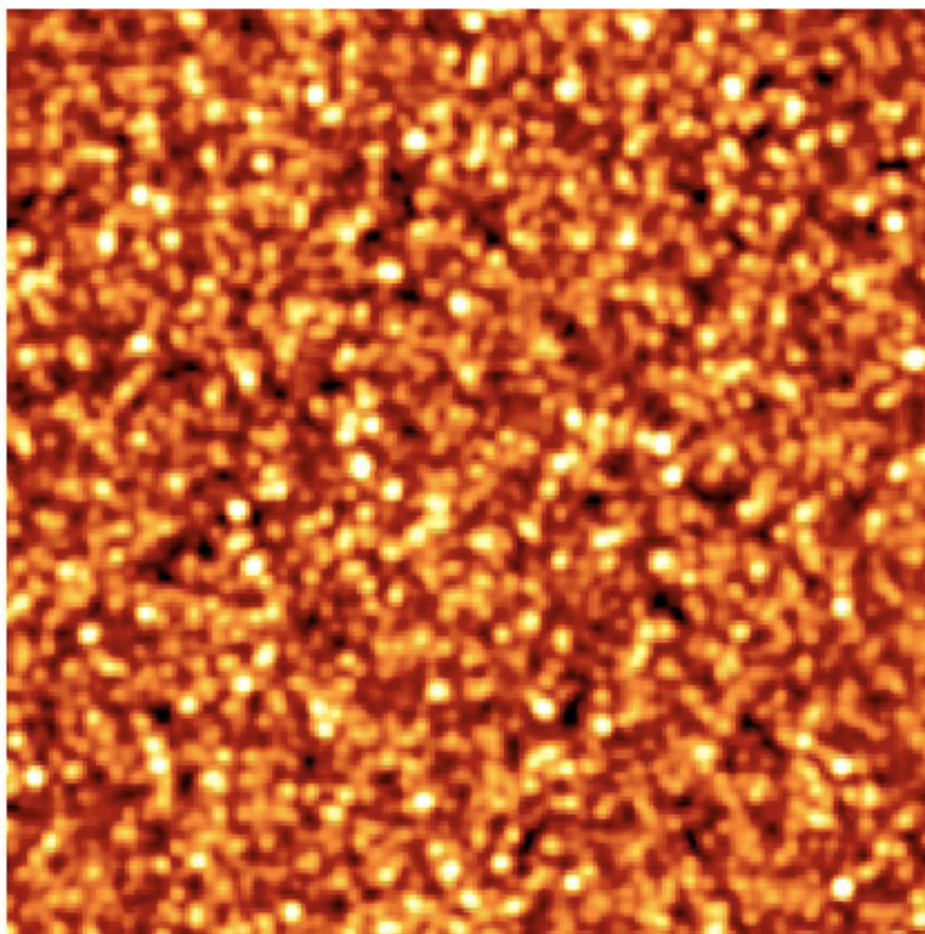
- Halo hosting a  $> 30 \text{ mJy}$  galaxy:  $M = 10^{12.6} M_{\text{sun}}$
- Satellites in more massive halos  $M \sim 10^{13.1} M_{\text{sun}}$



# Use Maps to Measure Clustering



$S > 20 \text{ mJy} : 1,200/\text{deg}^2$



$S < 20 \text{ mJy} : 480,000/\text{deg}^2$

# Large Scale Structure

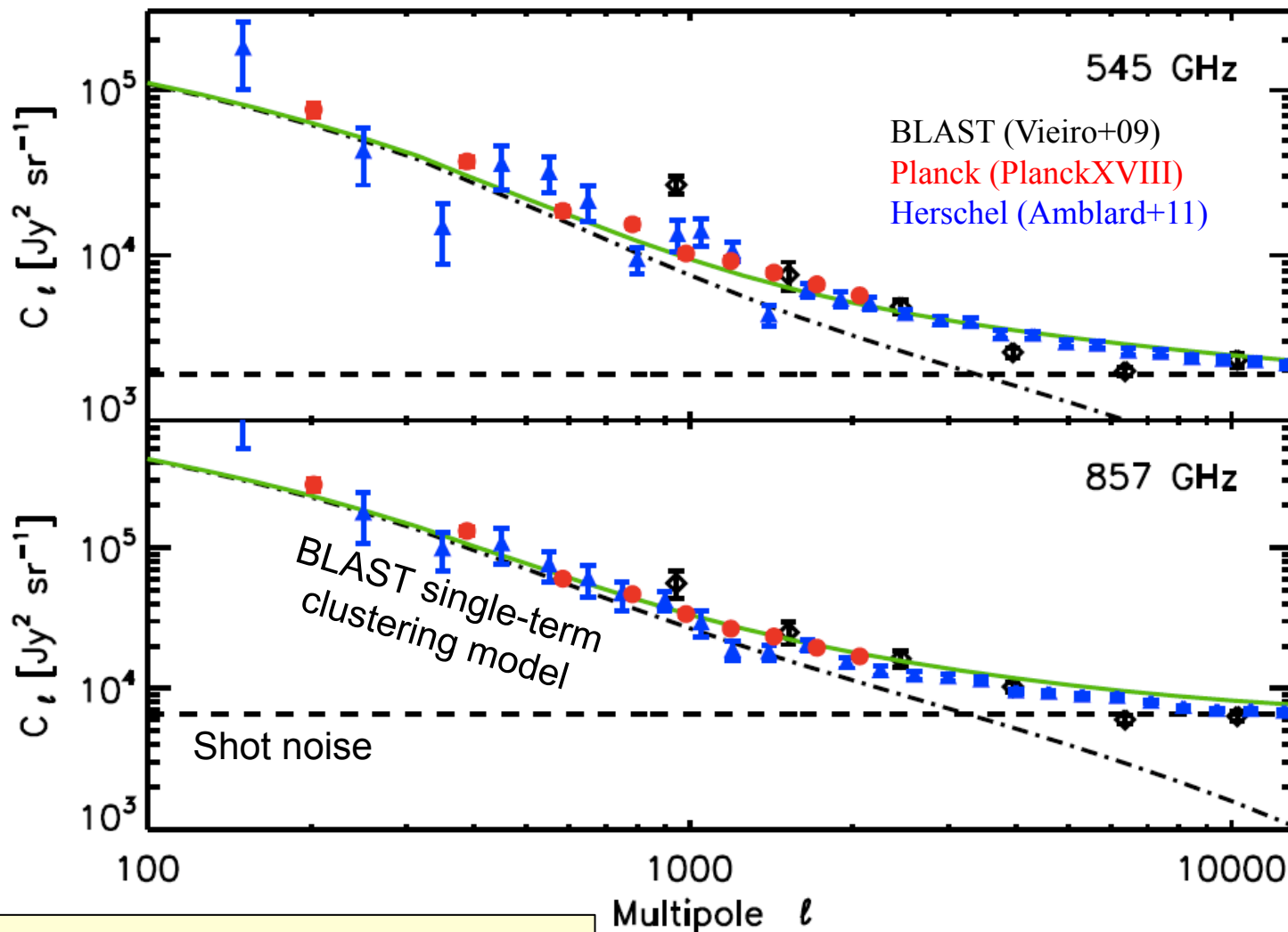
## Lockman Survey Field



3.6°

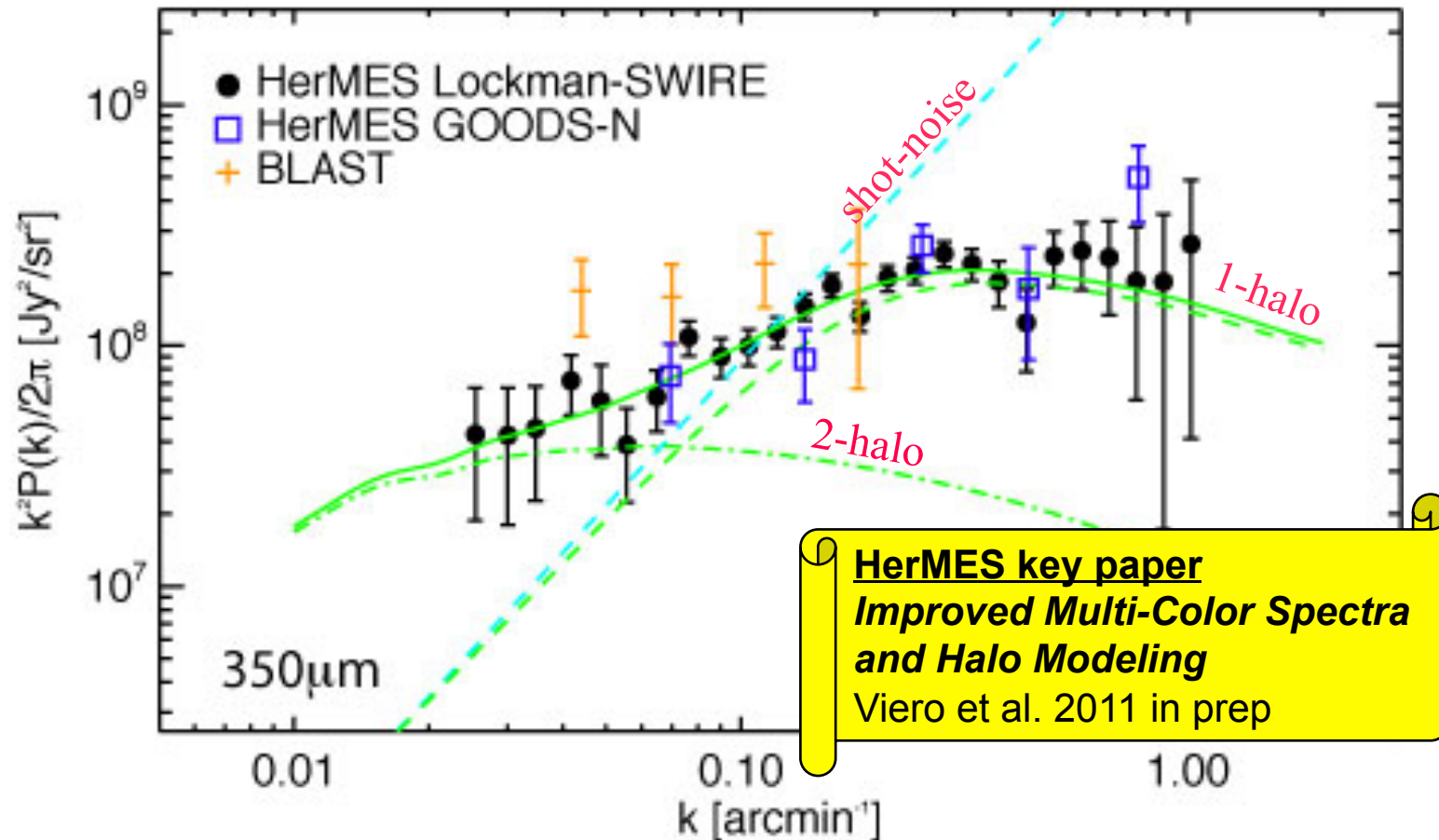


# Combined EBL Spatial Power Spectra



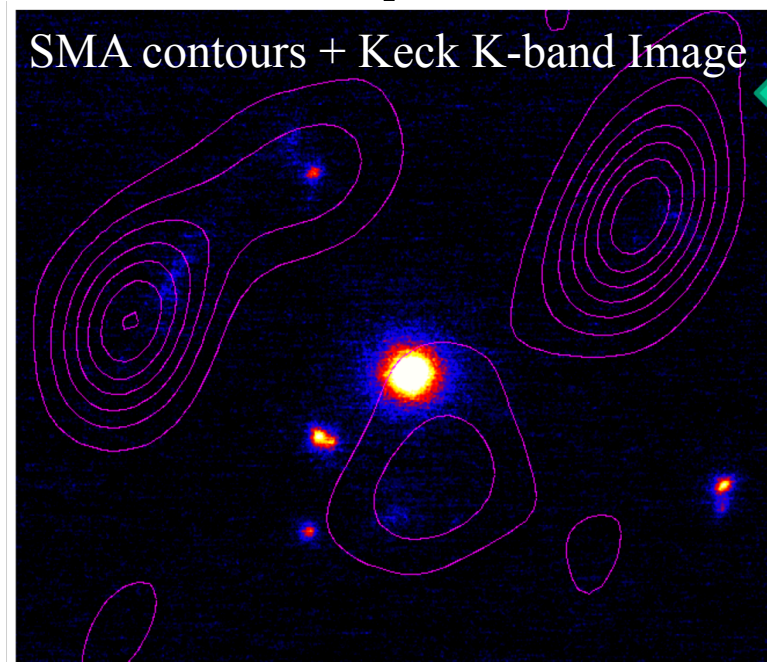
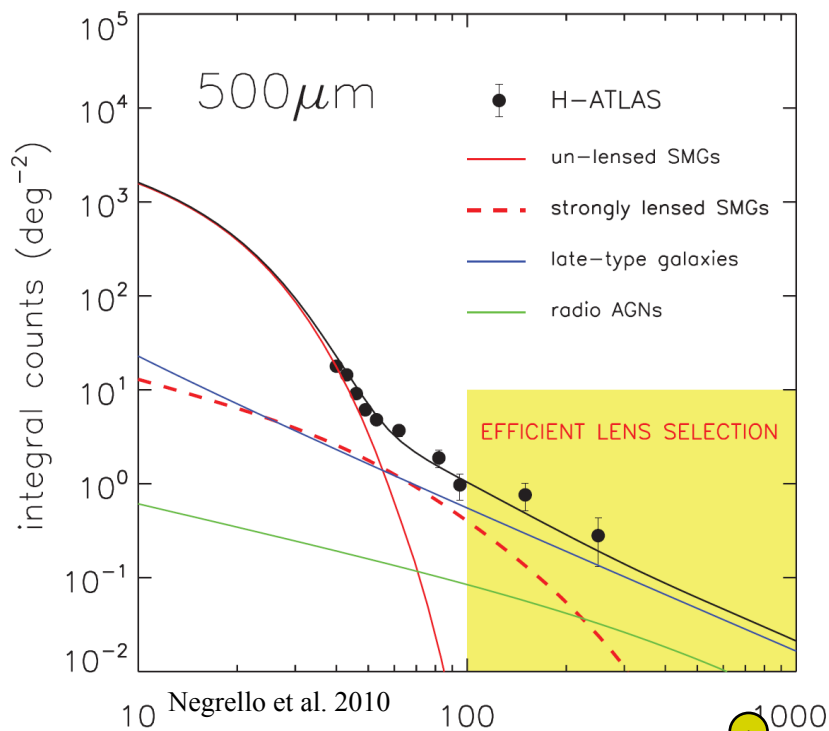
- SPIRE diffuse calibration corrected
- BLAST flux cut
- No cirrus removal

From Planck XVIII (arXiv 1101.2028)



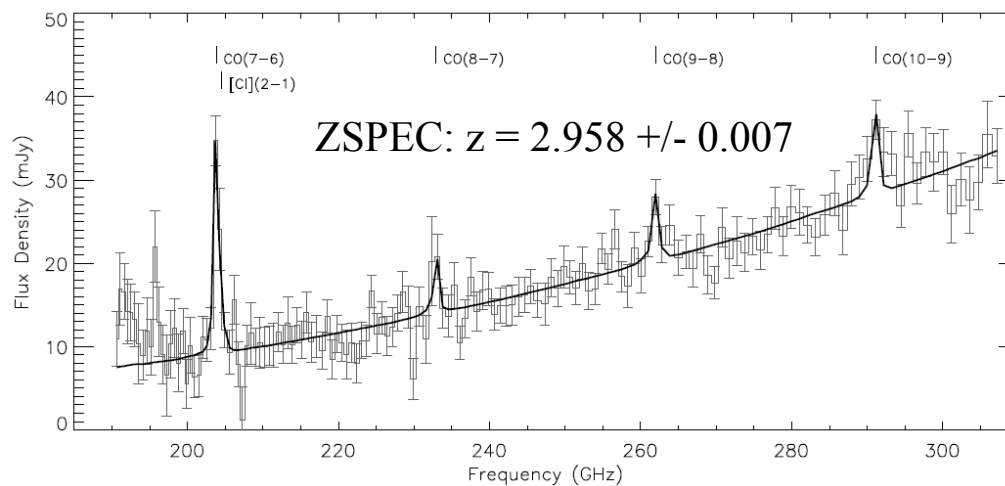
- Spectrum shown after removing shot noise and cirrus
- Clear 1-halo clustering seen
- Informs minimum halo mass  $\sim 3 \times 10^{11} M_{\text{sun}}$

Bright source found by HerMES: 420 mJy at 250 um



**HerMES key paper**  
**Lensing source statistics**  
 Wardlow et al. 2011 in prep  
 Talk by A. Cooray on Monday

Gavazzi et al. 2011, Scott et al. 2011,  
 Conley et al. 2011, Riechers et al. 2011

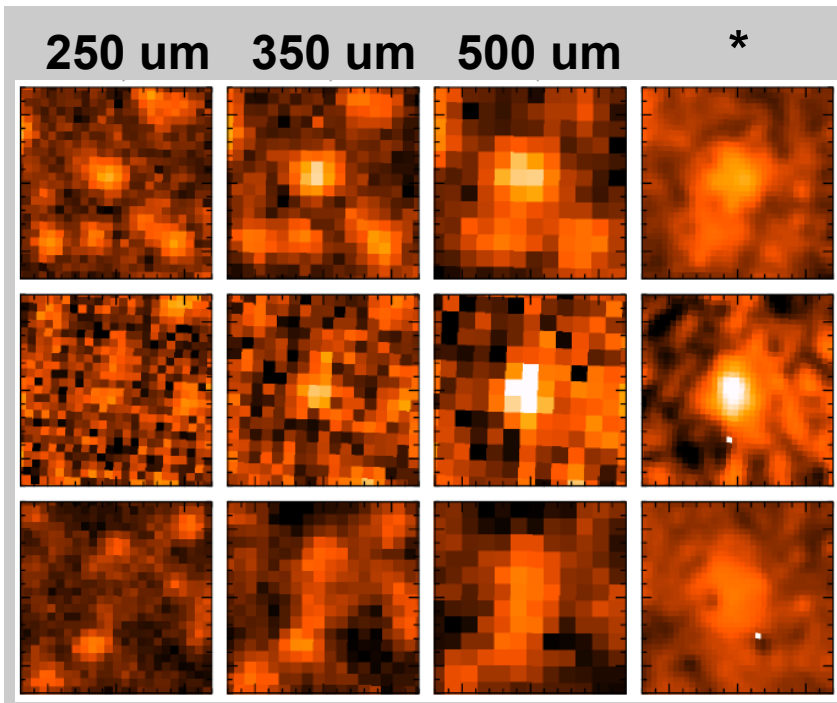




# High Redshift Galaxies in HerMES

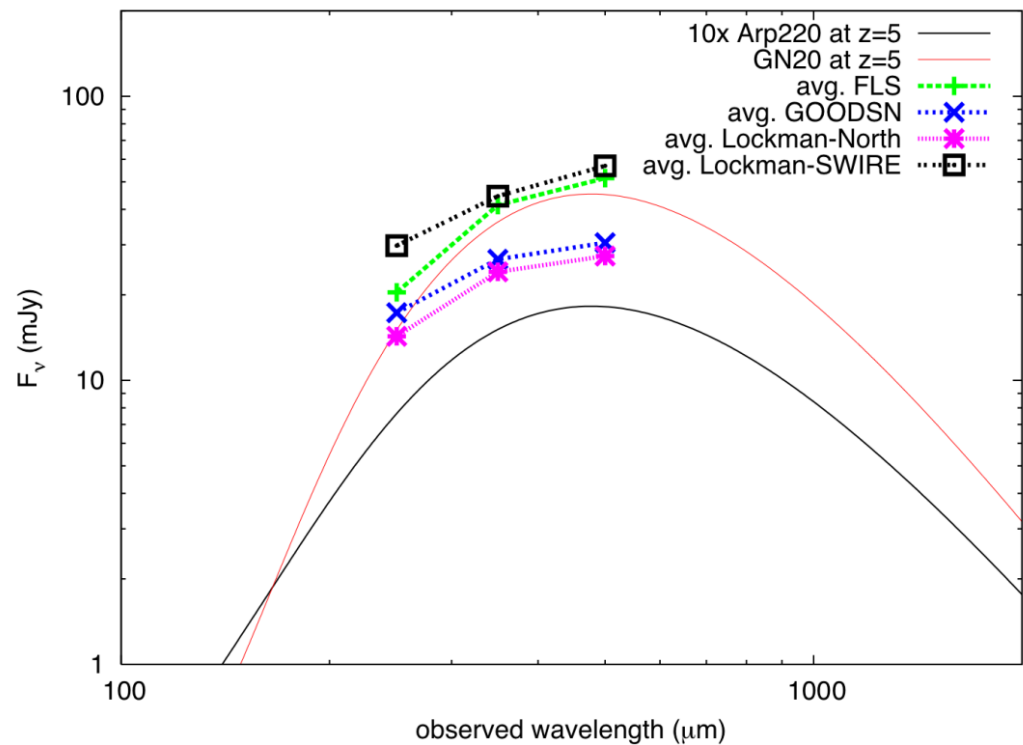
## 500 $\mu\text{m}$ peaked sources

Three examples:



\*Confusion reduced S(500) – fS(250)

Average spectra of sources detected in 4 HerMES fields compared to templates:



*These could be:*

*$z = 1.5, T_{dust} = 20 \text{ K ULIRGs or}$*

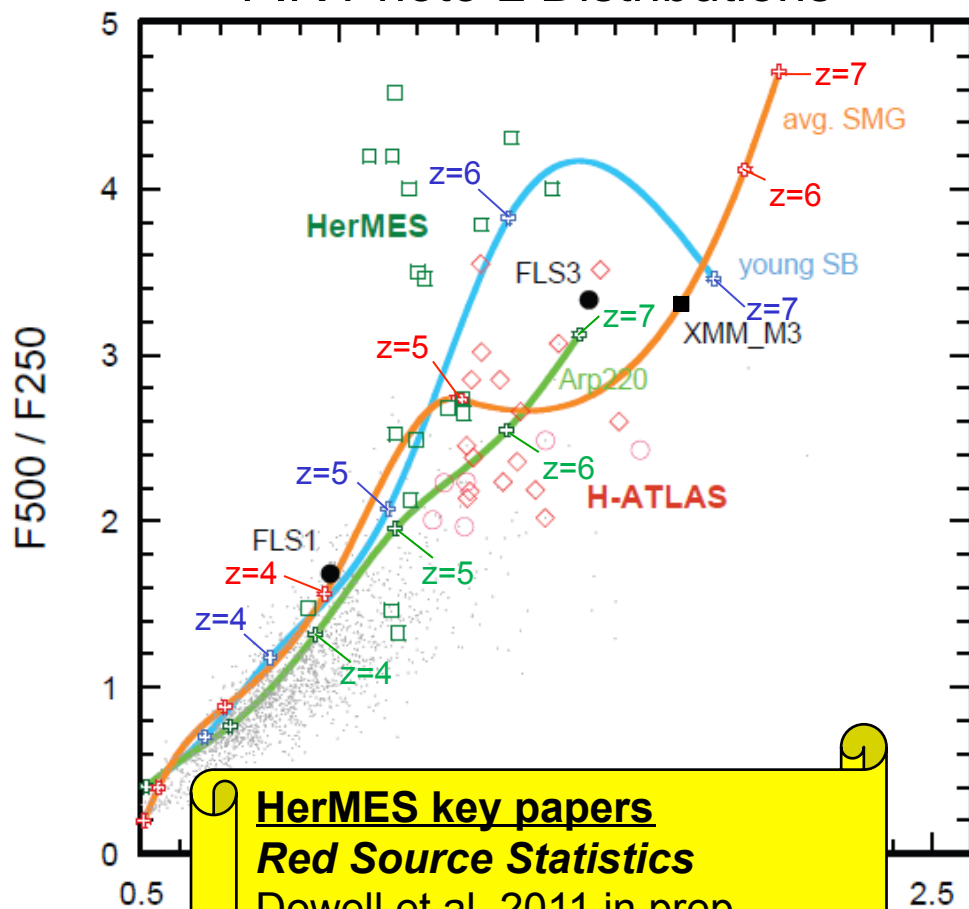
*$z = 5, T_{dust} = 45 \text{ K HLIRGs...}$*





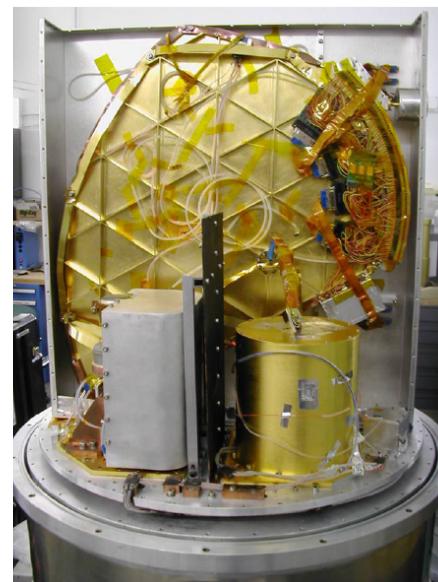
# High Redshift Candidates Follow-Up

## FIR Photo-z Distributions



**HerMES key papers**  
**Red Source Statistics**  
Dowell et al. 2011 in prep  
**High-z Galaxy Properties**  
Riechers et al. 2011 in prep

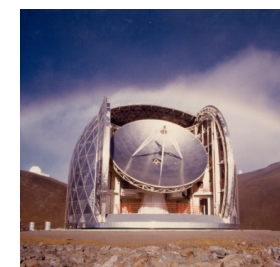
Obtain mm-wave redshifts (CO, C+)!!



ZSPEC: Bolometric spectrometer



APEX



CSO



PdBI



CARMA

# HerMES: What's Coming Next?

ER2 Data Release!!!

Catalogs Now Available <http://hedam.oamp.fr/HerMES/>

- A2218
- L2 ECDFS
- L3/5 Lockman Hole
- L6 Bootes
- L6 FDS
- L6 XMM-LSS

DR1 Release: Key Papers Conference

- + Joint PEP/HerMES Projects
- + Collaboration Projects
- + Follow up Observations

HeLMS Shallow Survey Approaching

DR2 Release at Mission End

**See HerMES Talks at this conference**

A. Cooray: Studies of Lensed Sources

D. Burgarella: HerMES Detections of Lyman-Break Galaxies

L. Wang: Halo Model of Galaxy Evolution

M. Bethermin: Constraints on the CIB

D. Farrah: Diagnostics of AGN Feedback